Third-harmonic-generation of a diode laser for quantum control of beryllium ions

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Abstract: We generate coherent ultraviolet radiation at 313 nm as the third harmonic of an external-cavity diode laser. We use this radiation for laser cooling of trapped beryllium atomic ions and sympathetic cooling of co-trapped beryllium-hydride molecular ions. An LBO crystal in an enhancement cavity generates the second harmonic, and a BBO crystal in a doubly resonant enhancement cavity mixes this second harmonic with the fundamental to produce the third harmonic. Each enhancement cavity is preceded by a tapered amplifier to increase the fundamental light. The 36-mW output power of this all-semiconductor-gain system will enable quantum control of the beryllium ions’ motion.

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References and links
1. Introduction

Trapped atomic ions have proven useful in a variety of fundamental and applied physics experiments. The beryllium ion (\(^{9}\text{Be}^{+}\)) in particular has been used for quantum information processing [1, 2], quantum simulation [3], quantum logic spectroscopy for optical atomic clocks [4, 5], as well as sympathetic cooling of molecules [6, 7] and highly charged ions [8]. The lowest transition in \(^{9}\text{Be}^{+}\) from the ground state to an excited electronic state is in the ultraviolet (UV) near 313 nm [9]. There are no convenient materials with optical gain at that wavelength, so \(^{9}\text{Be}^{+}\) laser systems require nonlinear frequency conversion of one or more laser sources. Such systems balance complexity, cost, and output power. Existing techniques include second-harmonic generation (SHG) of dye lasers [9], doubly resonant sum frequency generation (SGF) of a Ti:sapphire laser with a frequency-doubled Nd:YAG laser [10], fifth-harmonic generation of an amplified fiber laser [11], frequency-doubled light from SGF of two amplified fiber lasers [12, 13], fourth-
harmonic generation of an amplified diode laser, and SHG of diode lasers tuned by extreme temperatures [14]. The only published result using these temperature-tuned diode lasers achieved 35 µW of UV light [14], though outputs in the few-milliwatt range should be possible [15].

Our approach combines the versatility and low cost of a diode laser with the relatively higher powers available with tapered amplifiers in the infrared (IR). By use of two nonlinear crystals in power-enhancement cavities, we generate the third harmonic of 939 nm. Our system’s output power is 1000 times that of [14], though it comes with the added complexity of a doubly resonant SFG stage. Figure 1 provides a schematic of the system. An external-cavity diode laser (ECDL) serves as a tunable master laser at the fundamental frequency. The first nonlinear stage doubles the frequency in lithium triborate (LBO, LiB$_3$O$_5$). This second-harmonic light is summed with the fundamental in $\beta$-barium borate (BBO, $\beta$-BaB$_2$O$_4$) in a doubly resonant cavity. Our approach is similar to that in [16], except we seed with a diode laser instead of a Ti:Sapphire laser. The ability for rapid frequency modulation through the diode laser current allows us to use the doubling cavity itself to narrow the laser linewidth considerably and to lock the enhancement cavities with a phase-modulation (Pound–Drever–Hall, PDH) scheme [17] instead of a polarization (Hansch–Couillaud) scheme [18]. Our demonstrated output power is sufficient that this laser could be used for stimulated Raman transitions with several-gigahertz detunings [19, 20].

2. Master laser

The master laser is a home-built ECDL. The single-mode diode (Sheaumann M9-940-0200-D5P) is collimated with an aspheric lens. A diffraction grating (Thorlabs GH13-18U, 1800 lines/mm, efficiency $\approx$ 30 %) in the Littrow configuration provides feedback. The grating and collimation package are held and aligned with aluminum flexure mounts similar to those in [21]. The assembly is temperature-stabilized near room temperature. Feed-forward [22, 23] from the piezo voltage to the diode current by use of an analog circuit allows mode-hop-free tuning over at least 7 GHz. A typical output power of the master laser is 70 mW, which provides sufficient power for seeding two tapered amplifiers as well as monitoring and diagnostics. A bias tee allows radiofrequency modulation of the laser current [24, 25]. We apply a signal at 14.74 MHz and $-25$ dBm, which is used in stabilizing the length of the power-enhancement cavities.

We tune the laser near resonance with the $^9$Be$^+$ transition by use of a traveling Michelson interferometer [26]. This wavelength meter is referenced to a stabilized HeNe laser and measures the IR frequency to several parts in $10^7$.

We stabilize the laser against long-term drift by use of a scanning Fabry–Pérot cavity and a stabilized HeNe laser [27]. The cavity has a confocal design and a 750 MHz free-spectral range. Both the IR laser and the HeNe are incident on the cavity, and we monitor the transmission resonances of each while sweeping the cavity length. Dichroic mirrors combine and separate the two beams. After each sweep, a computer determines the locations of each resonance and feeds back to the ECDL grating and current to keep the IR wavelength stable relative to the HeNe wavelength. We enclose the cavity in a sealed tube to avoid pressure changes to the index of refraction of air [28]. One scan—feedback cycle takes approximately 40 ms. With this system, the long-term drift of the IR laser should track that of the stabilized HeNe. We have measured drifts of less than 200 kHz over three hours.

The short-term frequency noise of the ECDL is dominated by current noise. We measure its linewidth (as an overlapped Allan deviation) by monitoring the PDH signal of the SHG cavity (Sec. 3) on timescales faster than our piezo bandwidth. The free-running linewidth is 100 kHz at 15 µs, which is consistent with the current-noise specifications of our commercial controller (Thorlabs LDC202C). We use this PDH signal to feed back on the current and narrow the linewidth by a factor of five. Figure 2 displays the reduction of the frequency noise as a function of time. Independent measurements of the SFG cavity’s PDH signal confirm that the
frequency noise is indeed reduced. Narrowing the linewidth significantly reduces amplitude fluctuations in the second- and third-harmonic light.

Two tapered amplifiers (m2k Laser m2K-TA-0950-1500) boost the power available to the nonlinear conversion stages. Each amplifier produces 1.5 W of the fundamental light with an order of 10 mW of input light. A combination of aspheric, cylindrical, and spherical lenses couple this light into polarization-maintaining fibers. Including loss from optical isolators and imperfect matching of the TA mode to the fiber, we typically deliver 35–40% of this light to the nonlinear stages.

3. **Second harmonic generation (SHG)**

The second-harmonic generation (SHG) stage uses a nonlinear crystal in a power-enhancement cavity that is resonant with the fundamental. For this stage, we considered two crystals: LBO and bismuth triborate (BiBO, BiB₃O₆). The BiBO crystal benefits from a four-times larger effective nonlinear coefficient (3.36 pm/V in BiBO versus 0.81 pm/V in LBO). BiBO suffers
from a four-times larger walkoff, however. The most dramatic problem with BiBO is a significant photorefractive effect at high blue powers [29, 30]. While we eventually chose to run with LBO, we successfully generated blue light with BiBO in an enhancement cavity. We found that we could produce approximately 40 % more 470 nm light with BiBO than with LBO for the same input powers. At the highest powers, however, the photorefraction caused the output power to decline on timescales of tens of minutes and eventually induced oscillations in the crystal’s refractive index (and thus in the cavity mode, cavity input coupling, and blue conversion) with periods of several seconds. Only with 470 nm output powers below 40 mW would the SHG stage produce consistent output for several hours at a time. While the LBO crystal has a lower nonlinear coefficient, it displays no photorefraction at these powers.

The 10-mm-long LBO crystal is cut for type-I conversion (ooe) with critical birefringent phase matching at room temperature ($\theta = 90^\circ, \phi = 19.8^\circ$) [31]. The index of refraction is 1.608 for both wavelengths. Its input and output facets are cut at Brewster’s angle for the fundamental wavelength ($\theta_B = 58.1^\circ$). Compared to an anti-reflection coating, the Brewster-cut facet has lower loss at the fundamental and allows for compensation of astigmatism in the cavity due to the off-axis incidence on the spherical cavity mirrors. An important disadvantage is an approximately 20 % reflection loss of the second harmonic, which is polarized orthogonal to the fundamental. A side of the crystal is polished, and we mount the crystal such that the reflected light is available for monitoring and diagnostic purposes.

The cavity mirrors are positioned such that the primary waist (in the crystal) is first-order insensitive to the placement of the focusing mirrors [12] and astigmatism is compensated to create a circular secondary waist. The round secondary waist simplifies mode-matching into the cavity. The curved cavity mirrors (M3, M4) have a 75-mm radius of curvature and are placed 39 mm from the crystal facet with the full-opening angle of the incident-to-reflected beam equal to $23.6^\circ$. We calculate that the focus of the fundamental in the crystal has a horizontal waist of 48.6 $\mu$m and a vertical waist of 31.7 $\mu$m. The cavity free spectral range is measured to be 488 MHz.

Double-refraction in the crystal causes second-harmonic walkoff of $\rho = 11.5$ mrad. This walkoff, combined with our crystal length $l$ and the wavenumber $k_1$ of the fundamental beam, give a crystal parameter (in the notation of [32]) $B \equiv \rho (lk_1)^{1/2}/2 = 1.9$. Both the Boyd–Kleinman analysis [32] and that of Freegarde, et al. [33], which extends the analysis to elliptical beams, indicate that our focus is somewhat larger than optimal. It should be possible to increase the blue output at the expense of either a significantly longer cavity or an elliptical secondary waist.
Fig. 3. Output power versus input power for the SHG cavity. The theory curve corresponds to nonlinear coefficient $\gamma_{\text{SHG}} = 1.0 \times 10^{-5} \text{ W}^{-1}$, passive cavity loss of 0.30 %, and includes a 20 % reflection loss at the crystal output facet.

The positions and angles of the mirror mounts are fixed with dowel pins in a custom-machined aluminum baseplate. One flat mirror (M2) is glued to a piezoelectric transducer, which allows active stabilization of the cavity length. This mirror is small (6.0 mm diameter, 2.3 mm thick) to minimize mass-loading the piezo. We monitor the reflected light from the cavity and lock it to resonance with the Pound–Drever–Hall technique [17]. The cavity’s finesse of 600 is high enough that the laser’s modulation sidebands are nearly completely reflected when the cavity is on resonance. The ideal input-coupler (M1) has a transmission at the fundamental that matches the loss in the cavity (including conversion loss) [34]. We have tried mirrors with transmissions of 3.0 %, 2.2 %, 0.9 %, and 0.4 % and found the latter two to give comparable conversion efficiency. The data here were taken with the 0.4 % (99.6 % reflection) input coupler. Two spherical singlet lenses mode-match the collimated fiber output to the cavity.

Figure 3 shows the output power of the SHG stage as a function of input power. Also plotted is a model [34] that corresponds to a second-harmonic generation coefficient of $\gamma_{\text{SHG}} = P_{2\omega}/P_{c}^{2} = 1.0 \times 10^{-5} \text{ W}^{-1}$ and a passive cavity loss (excluding loss from conversion and the input coupler transmission, and stated as an equivalent reflectivity) of $R_{m} = 0.9970$. Approximately 15 % of the incident fundamental light is reflected from the input coupler in a manner more consistent with a mode mismatch than an impedance mismatch. At 600 mW of IR light incident on the cavity, we measure 120 mW of blue light through mirror M4. The transmission through the input coupler is 81 %. The IR transmission through the high-reflectivity mirror M3 (transmission 0.04 %) is 47 mW, which indicates a circulating power of $P_{c} = 117$ W. This power-enhancement of nearly 200 agrees with that calculated from the cavity linewidth when scanning the cavity length.

4. Sum frequency generation (SFG)

The second harmonic light mixes with additional fundamental light in the sum-frequency generation (SFG) stage. This stage is doubly resonant to enhance the power at both wavelengths. The 10-mm-long BBO crystal is cut for type-I conversion (ooe) with critical birefringent phase matching at room temperature ($\theta = 36.0^\circ$) [31]. The crystal is cut for normal incidence and anti-reflection coated at the fundamental and both harmonics. Since we want the fundamental and second harmonic to traverse nearly identical paths in the cavity, a Brewster cut will not work here. Dispersion in the BBO means the two beams have slightly different indices of refraction and thus would not be overlapped after refracting at the crystal facet. The indices are 1.657,
Fig. 4. Model of output versus input powers for the SFG cavity. Measured output powers agree to within 15% of the model. Input powers accessible with our system are up to approximately 630 mW in the fundamental (1\(\omega\)) and 120 mW in the second harmonic (2\(\omega\)). The model parameters are a nonlinear coefficient \(\gamma_{\text{SFG}} = 2.94 \times 10^{-4} \text{ W}^{-1}\), passive cavity losses of 1.9% in the fundamental and 1.8% in the second harmonic, and mode-match efficiencies of 79.4% in the IR and 84.7% in the blue.

1.681, 1.673 for 940 nm, 470 nm, and 313 nm.

The dispersion in the BBO crystal must be compensated by other dispersive elements. We use a pair of 1.0-mm-thick fused silica plates [16]. In addition to dispersion compensation, the second plate compensates any displacement from the first plate. Each plate is mounted near Brewster’s angle. Rotating the plates through 10 mrad provides a half-wavelength shift at 470 nm in the relative path length. Drifts in the BBO dispersion from temperature or other effects require active control of the plate angles. We mount them on counter-rotating galvanometers (Cambridge Technology model G108). It is important that the galvos be of the “open-loop” design. While “closed-loop” galvos dominate the current market because of their faster response times, this purpose does not need the increased speed and we have found their encoders to have insufficient resolution for this task.

Because of the normal-incidence crystal cut, the astigmatism from the off-axis incidence on the spherical mirrors cannot be compensated. Thus we make the cavity opening angle as small as practical—a full-angle opening of 12°. We again position the cavity mirrors such that the primary waist is first-order insensitive to the placement of the focusing mirrors. These mirrors (M7, M8) have 75-mm radii of curvature and are placed 37 mm from the crystal facet. The secondary waist has an ellipticity of 0.91, which does not pose a significant challenge to mode-match. The primary’s waist is close to round with a calculated size of 27.3 \(\mu\)m for the fundamental. The second-harmonic’s waist is smaller by the square root of the wavelength ratio. The cavity free spectral range is measured to be 406 MHz.

The analysis for SFG is formally equivalent to that of SHG when the parameters are appropriately redefined and the beams have the same confocal parameter [32]. This is the case for our beams, which share the same optical resonator. The third-harmonic walkoff is \(\rho = 79.1\) mrad, which corresponds to a crystal parameter \(B = 16\). The resulting analysis indicates that our waist is nearly optimal (9% too large).

A dichroic mirror combines the fundamental and second harmonic and sends them towards the cavity. They couple in through the same mirror (M5), which is coated with a nominal 98%
Fig. 5. Performance of the third-harmonic (UV) output power on (a) short and (b) long timescales. The short-time performance is presented as the overlapped Allan deviation, which shows power fluctuations stay below 3% on most timescales. The long-time trace shows variation under 10%. Near the six-hour mark, the system automatically relocked in less than 10 s following a disturbance.

reflectivity at both wavelengths. The fundamental light comes from a single-mode, polarization-maintaining fiber and is mode-matched to the cavity by use of two spherical singlets. The second harmonic light comes from the SHG cavity and requires more conditioning. A spherical lens collimates the SHG output beam and two cylindrical lenses correct for residual astigmatism and ellipticity. Then two spherical lenses mode-match into the SFG cavity. In principle, fewer lenses are needed, but we find the alignment easier when the function of each lens is decoupled from the others. Along the way, a half-waveplate rotates the polarization by 90°.

The cavity is kept on resonance with two servos acting on different timescales. The light reflected off the input coupler is split by a dichroic mirror and monitored on two detectors. The reflected IR light is used in a PDH lock and controls a piezo just as in the SHG cavity. With the cavity actively locked to the IR light, the galvo angles are modulated by ≈ 10 µrad at 200 Hz. The signal from the reflected blue light is demodulated in a lock-in amplifier. A servo controller locks the dispersion such that the cavity is resonant in the blue as well.

To characterize the SFG conversion efficiency, we measure the UV output at several blue and IR input powers, while monitoring the circulating power and input-coupling fraction for each beam. We fit a model [35] of doubly resonant SFG to these data and plot the result in Fig. 4. Measured output powers agree to within 15% of the model. The best-fit model parameters are a sum-frequency generation coefficient \( \gamma_{\text{SFG}} = P_{3\omega}/(P_{1\omega,c} P_{2\omega,c}) = 2.94 \times 10^{-4} \text{ W}^{-1} \), passive losses of \( R_m = 0.981 \) and 0.982, and mode-match efficiencies of 0.794 in the IR and 0.847 in the blue. We measure 32 mW of UV output with 632 mW of IR and 105 mW of blue input. Leakage through mirror M7 (transmission 6.0 \times 10^{-5} at 940 nm and 8.1 \times 10^{-5} at 470 nm) indicate circulating powers of 27 W in the IR and 3.7 W in the blue. We have observed UV powers as high as 36 mW.

When running, the laser system uses five servos. The HeNe-stability-transfer lock is slow enough that we implement it in software. We implement the remaining four servos in two
commercial FPGA-based devices (Toptica DigiLock 110), which contain two servos each. Each
device controls one cavity. For the SHG cavity, one servo locks the cavity length with the piezo,
while the other servo feeds back to the master laser to remove short-term frequency noise. For
the SFG cavity, one servo locks the cavity length with the piezo, while the other implements the
galvo dispersion lock. In both cases, the two servos control processes on significantly different
timescales. The DigiLock includes an FPGA-based lock-in amplifier for the galvo lock. The PDH
locks use external mixers and amplifiers to condition the signal. The locks are interdependent and
need to work simultaneously to produce UV light. It is useful for the servos to have an automated
relock function to recover from sharp acoustic or vibrational disturbances. The DigiLock has
such a function, as do others including the open-source servo in [36].

The UV power is low-noise and stable. Power variations on short timescales are mostly below
3%. At intermediate times (milliseconds), fluctuations arise primarily from feedthrough of the
200 Hz galvo modulation. Figure 5(a) shows the overlapped Allan deviation from 1 µs to 1 s,
where the dominant features are oscillations from this modulation. The output power is stable to
better than 10% over long times. The system can stay locked for hours when left undisturbed.
Figure 5(b) shows such a multi-hour run with no user intervention. During the time shown,
the 470 nm second-harmonic power changed by less than one percent. Near the six-hour mark,
a disturbance required the servos to relock, which took approximately eight seconds. Similar
to others [16], we see no evidence of degradation of crystals or optical coatings from the UV
radiation. We do find that the second-harmonic output is much more stable when the cavity is
enclosed in a box with a low flow of dry nitrogen through it. Without this continuous flow, the
second-harmonic power drops in steps. The power can be recovered by translating or cleaning
the LBO crystal, which suggests the high optical powers in the SHG cavity may be entraining
dust or aerosols.

5. Use with trapped $^9$Be$^+$ ions

We use the laser to Doppler cool trapped $^9$Be$^+$ ions. We have used these atomic ions to sym-
pathetically cool BeH$^+$ molecular ions. The ions are confined in a linear radiofrequency (Paul)
trap [37]. The trap comprises machined OFHC copper electrodes mounted with titanium dowels
and screws onto MACOR spacers. The radial ion–electrode distance is $r_0 = 1.2$ mm. Two
opposing electrodes are electrically connected to a helical resonator [38]. Their electric potential
is driven at $\Omega/(2\pi) = 35$ MHz and several hundred volts. The other two electrodes are held at rf
ground. They are each divided into five segments with independently controllable potentials. The
center three electrodes have widths of $2z_0 = 3.0$ mm and inter-electrode gaps of approximately
0.1 mm. Potential differences of 4 V across the central three electrodes produce a $^9$Be$^+$ axial
frequency of $\omega_z/(2\pi) = 375$ kHz.

To load $^9$Be$^+$, we run a current through a beryllium-wrapped tungsten wire, which creates
a flux of neutral Be through the trap. An electron beam bombards the atoms, and any atom ionized within the trapping volume is trapped. We load BeH\(^{+}\) through a chemical reaction between trapped \(^9\)Be\(^{+}\) and H\(_2\) background gas. Our vacuum system has a precision leak valve to allow the controlled admission of gas, though the BeH\(^{+}\) loaded to date have been from less-controlled reactions with residual gas in our chamber. We tune the laser system to several hundred megahertz below the \(^9\)Be\(^{+}\) \(^2\)S\(_{1/2}^\rightarrow^2\)P\(_{3/2}^\) transition (313.13292 nm [9]). A Glan-laser polarizer and quarter-waveplate create a pure circular polarization to address the atomic ions’ cycling transition. Figure 6 shows the fluorescence of seven \(^9\)Be\(^{+}\) ions with gaps in the chain from two BeH\(^{+}\) ions.

6. Conclusion

In conclusion we have constructed a tunable UV laser near 313 nm capable of achieving 36 mW output power and based on third-harmonic-generation of an external-cavity diode laser and tapered amplifiers. We have demonstrated its use in optical-pumping, detection, and Doppler-cooling of \(^9\)Be\(^{+}\) atomic ions. This laser system produces sufficient power to drive stimulated Raman transitions [1] in \(^9\)Be\(^{+}\). By modulating the laser to add an optical sideband on-resonance with the \(^9\)Be\(^{+}\) transition, it should be possible to use this single laser for both coherent and dissipative control of \(^9\)Be\(^{+}\), including motional ground-state cooling and coherent sideband transitions. Such single-laser systems (based on dye or fiber lasers) have been implemented for the Mg\(^{+}\) ion [19, 20] by use of an electro-optic modulator to add the sideband before a SHG stage. In our setup, we could use such a modulator or could modulate the current of the ECDL directly. We have been able to modulate our ECDL at frequencies as high as 13 GHz. In order for these sidebands to persist onto the UV output, the modulation frequency must be a multiple of the SFG cavity’s free-spectral range. Such a diode-laser-based system will prove fruitful for work with trapped \(^9\)Be\(^{+}\) ions in quantum information processing and precision measurements.

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